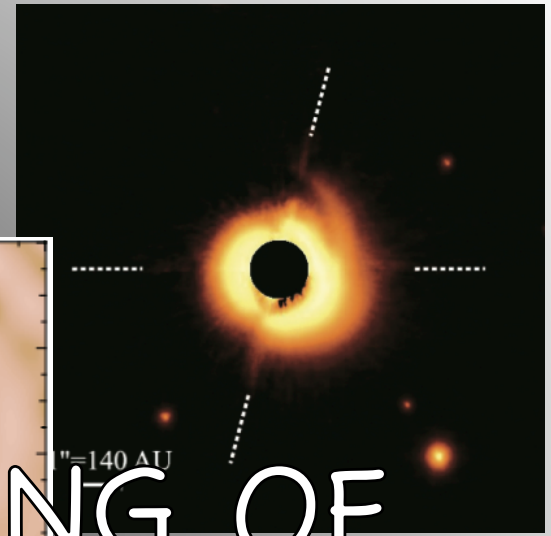
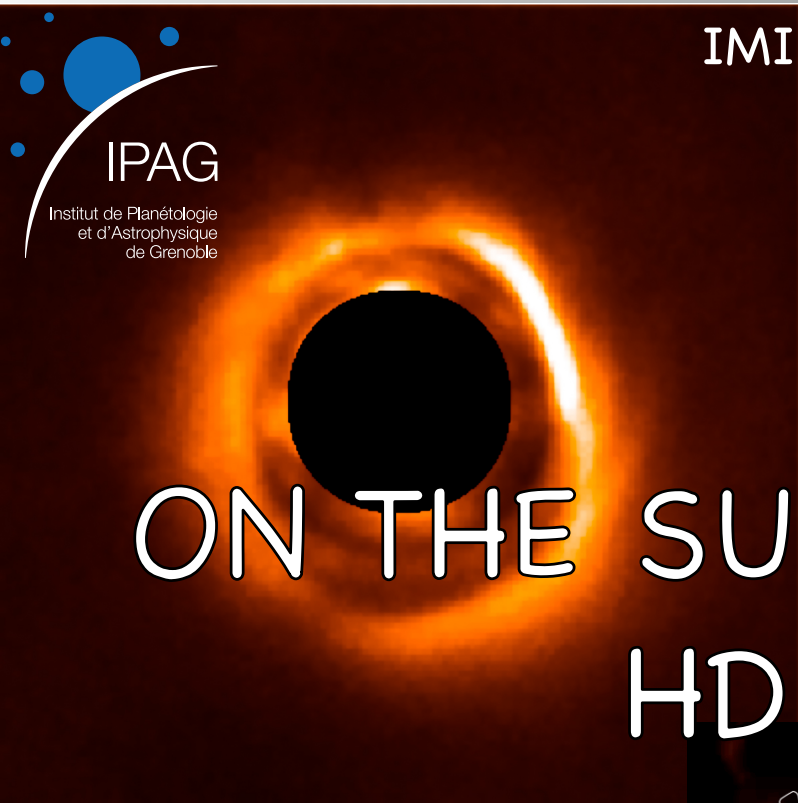


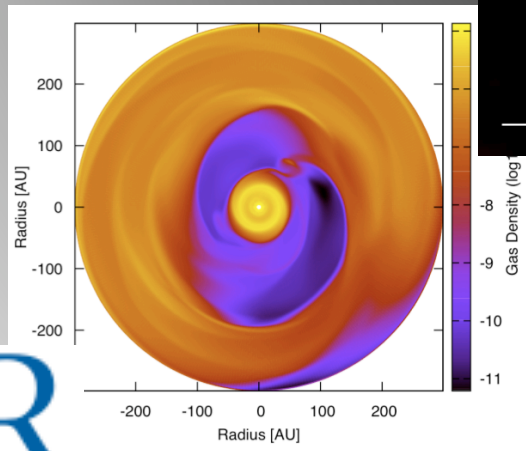
IMI 2012 - October



# ON THE SURROUNDING OF HD 142527



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# From the beginning up to now

## OUTLINE

- I. First high-contrast IR image
- II. Modelling the geometry
- III. A possible companion ?
- IV. Spirals, asymmetries and gaps
- V. An embedded forming planet ?

# Into the past HD 142527

Herbig Ae

F6IIIe

$d = 145 \pm 15$  pc *De Zeeuw et al. (1999)*

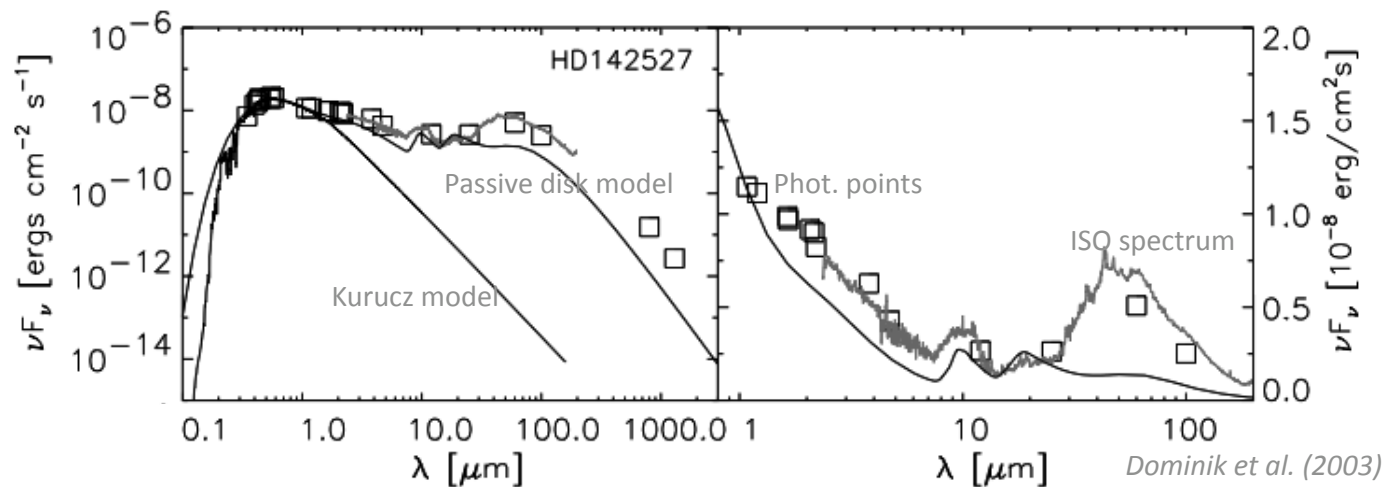
V mag = 8.34

K mag = 4.98

Age = 2 – 20 Myr

Member of Up. Cen. Lup. *De Zeeuw et al. (1999)*

$F_{\text{IR}}/F_{\star} = 0.92$  *Dominik et al. (2003)*



# I. First HC-IR image

*Fukagawa et al. (2006)*

Subaru / CIAO at K & H  
Coronagraphic & PSF reference star

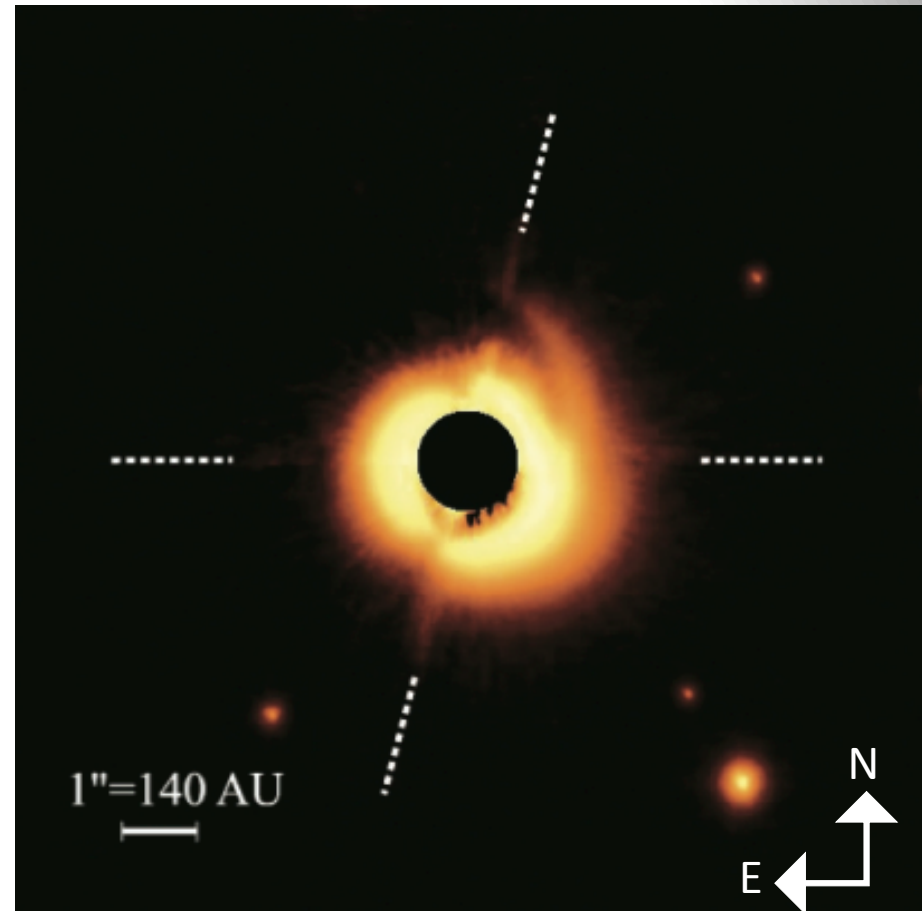
Offset of the center

Outer disk up to  $\sim 500$  AU

'Banana split'

Brightness asymmetries

Central binary or recent stellar encounter ?



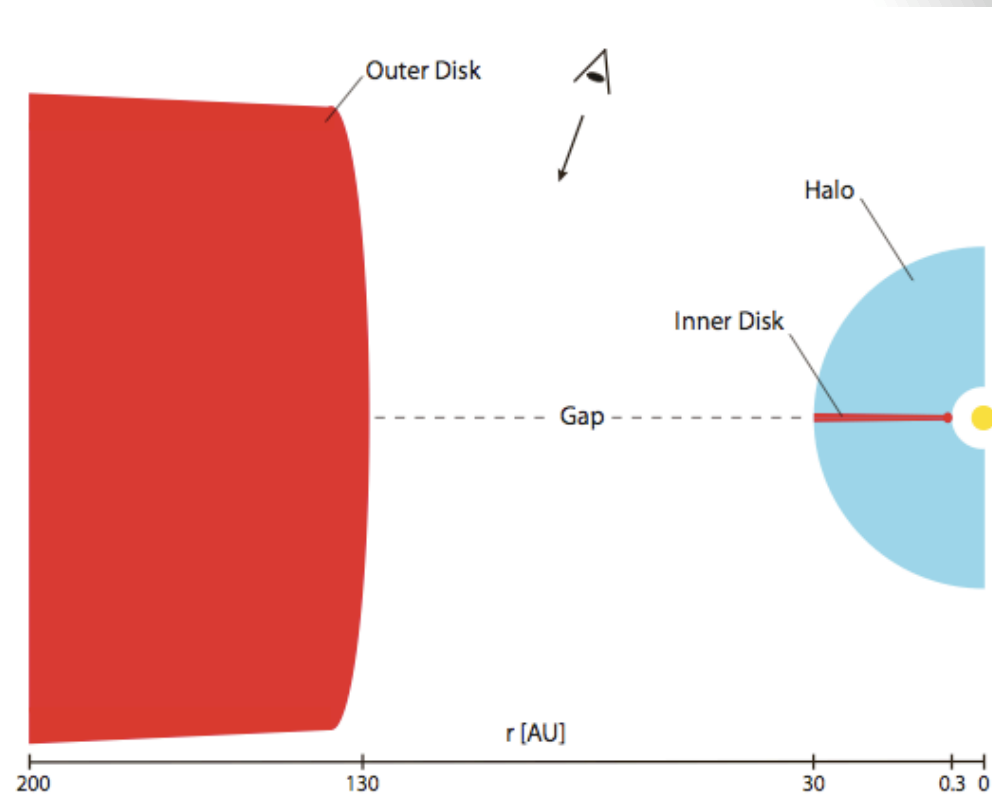
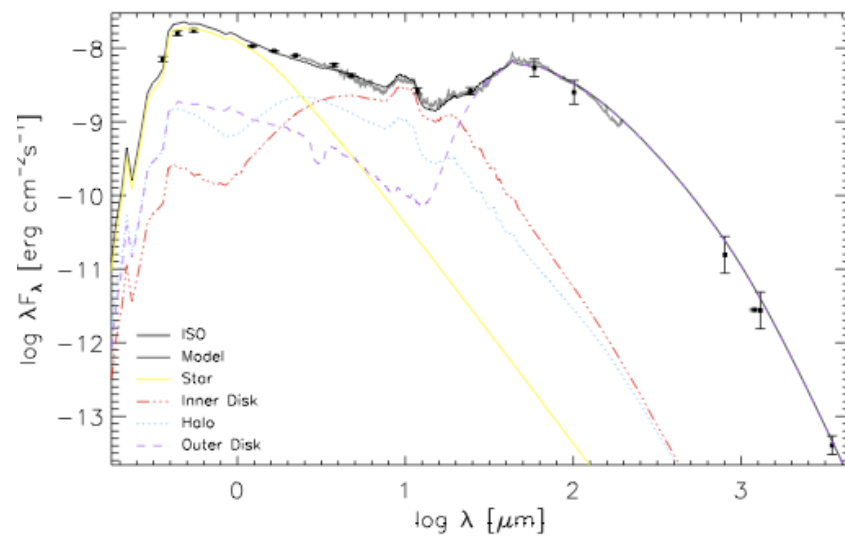
*Fukagawa et al. (2006)*

# II. Modelling the geometry

*Verhoeff et al. (2011)*

Compilation of observations [mm => nm]

3 geometrically distinct components : inner disk (thick), halo (thin) & outer disk (thick)



*Verhoeff et al. (2011)*

# III. A possible companion ?

NACO/SAM at H, K & L'

Fit the closure phase with a binary model

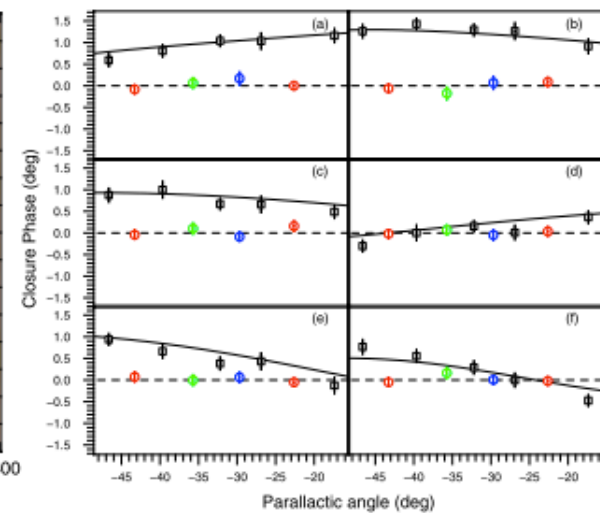
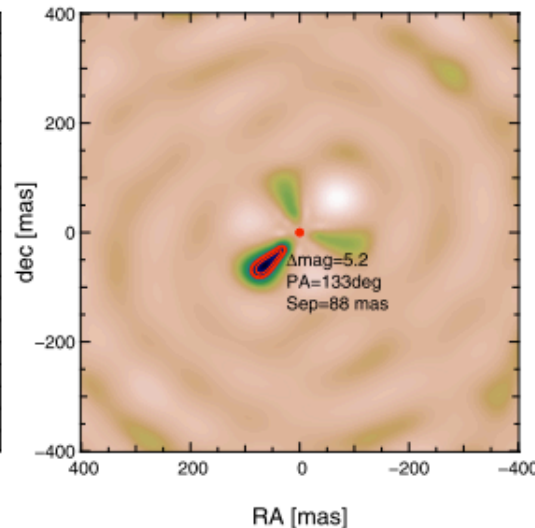
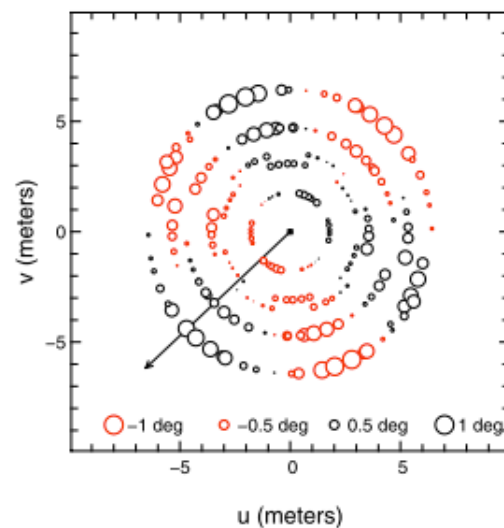
$\Delta L' = 5-6.5$  mag

$M \sim 0.1-0.4 M_{\odot}$  at  $12.8 \pm 1.5$  AU

Need confirmation !

*Biller et al. (2012)*

*MPIA/IPAG collaboration*

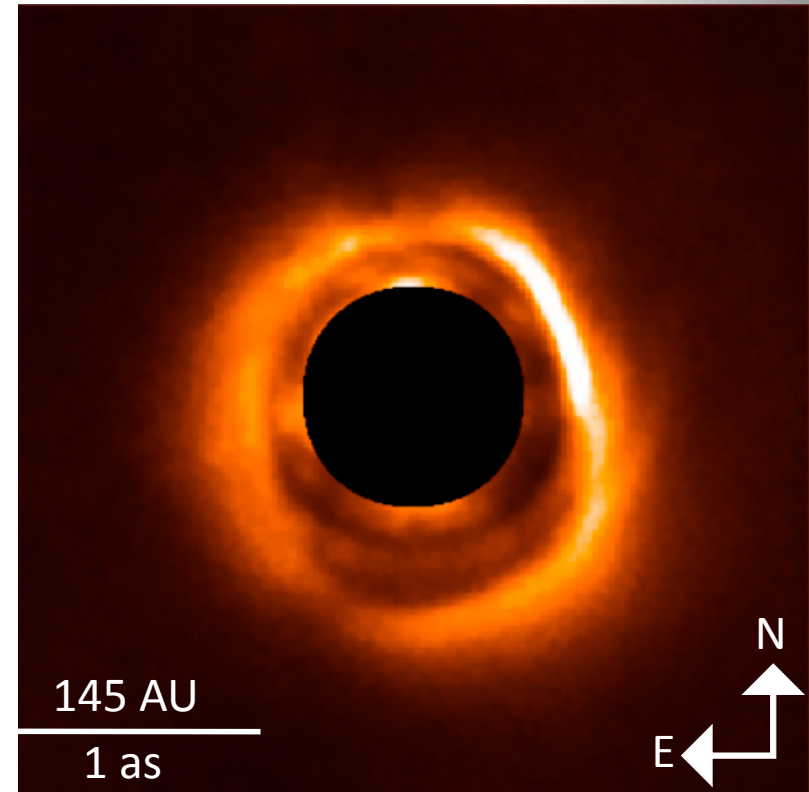


*Biller et al. (2012)*

# IV. Gaps, spirals and asymmetries

*Rameau et al. (2012)*

VLT/NACO at L'  
ADI & PSF reference subtraction



*Rameau et al. (2012)*



*Casassus et al. (2012)*

*Casassus et al. (2012)*

GEMINI/NICI at H, Ks & L'  
PSF reference subtraction

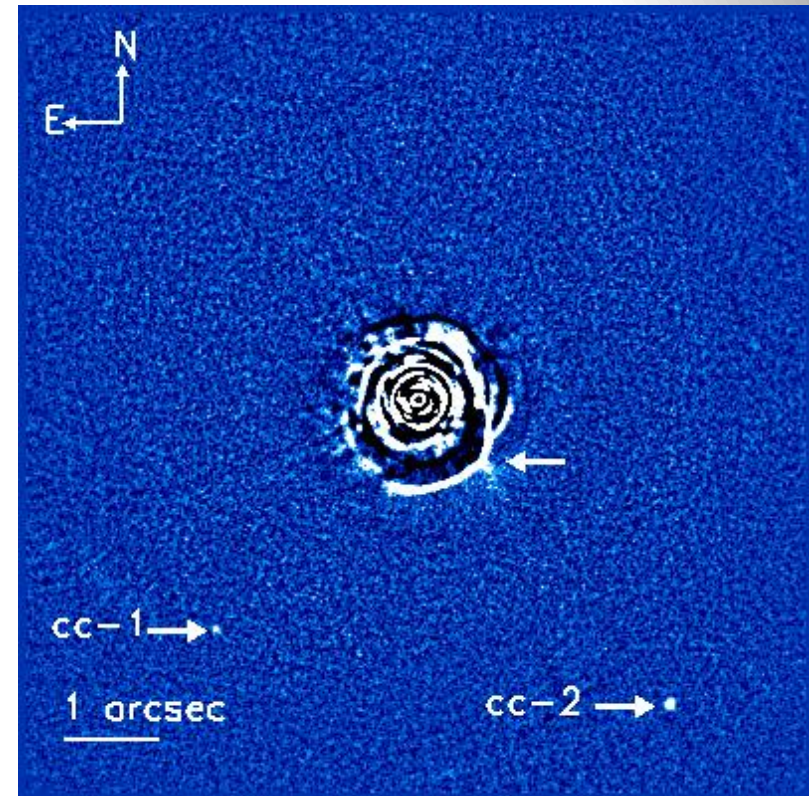


# IV. Gaps, spirals and asymmetries

ADI processing

Artifacts (Milli et al. 2012)

2 background sources



Rameau et al. (2012)

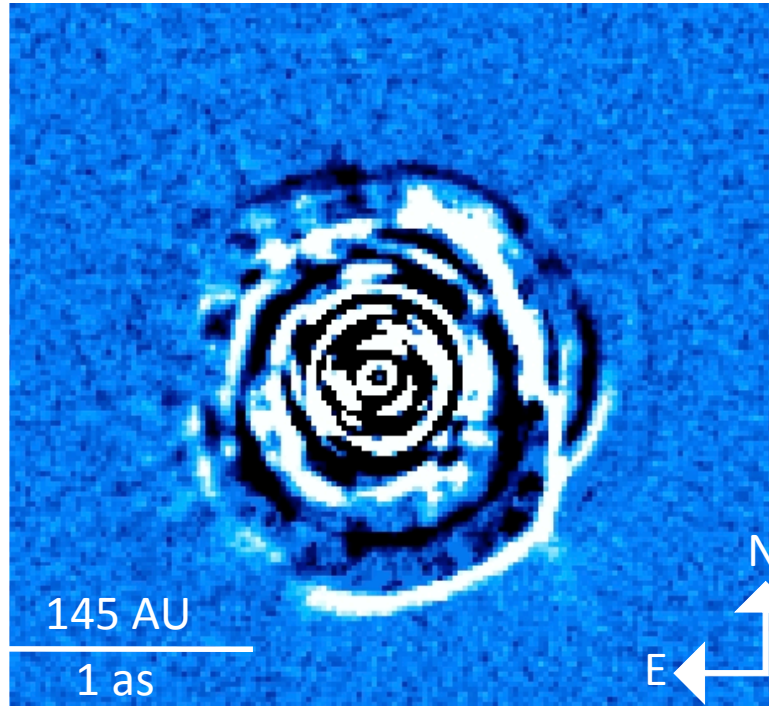


# IV. Gaps, spirals and asymmetries

ADI processing

Artifacts *( Milli et al. 2012 )*

2 background sources



*Rameau et al. (2012)*

# V. An embedded forming planet ?

Scenario of planet formation

*Casassus et al. (2012)*

FARGO (*Masset 2000*) simulation

10  $M_J$  at 90 AU

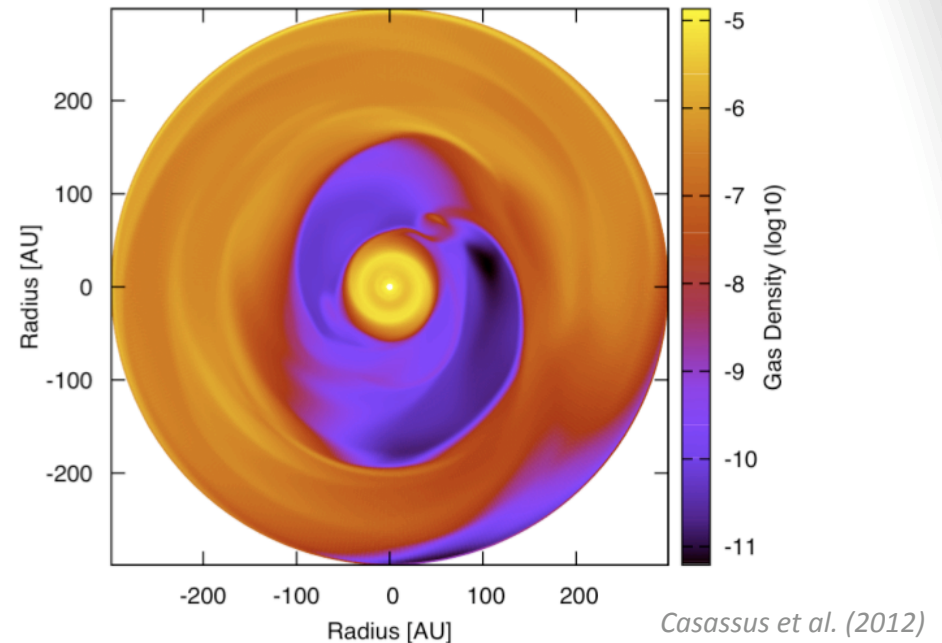
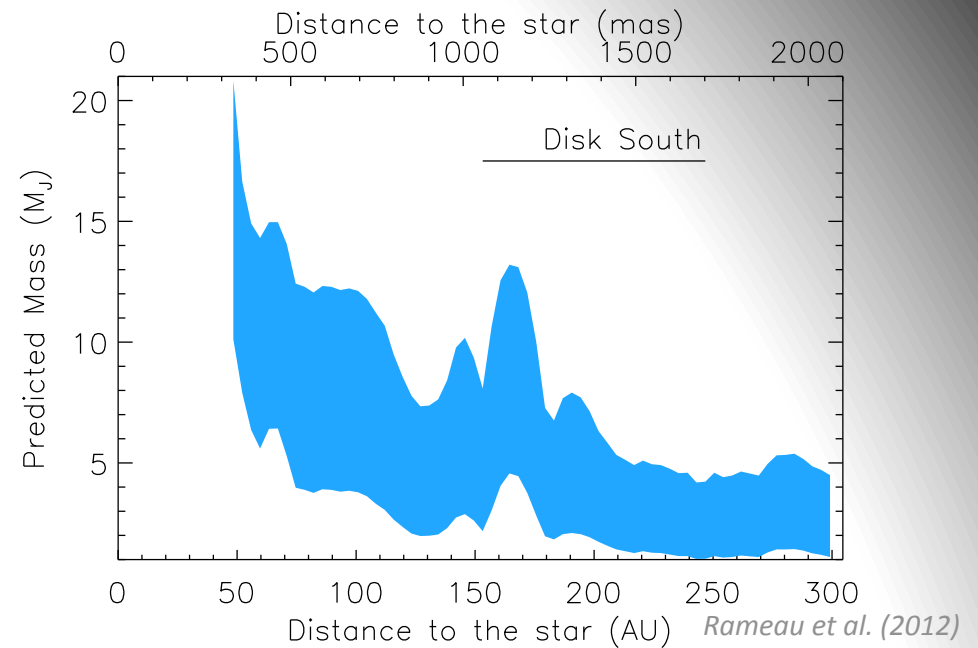
*Rameau et al. (2012)*

=> Rule out by our detection limits

Spirals : OK

Asymmetric gap : OK

Extended inner disk : No



# Going farther...

Need characterizations :

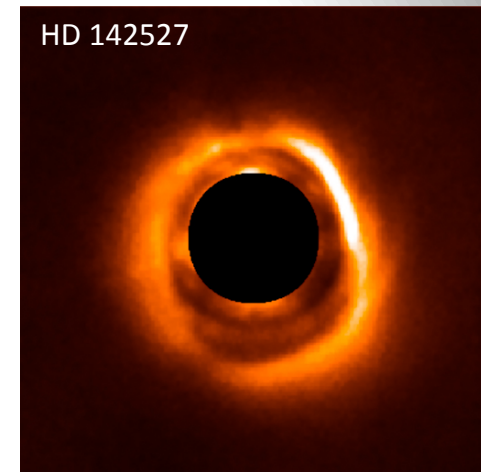
- Inner disk (VLT/PIONIER, SPHERE)
- Gas mass and distribution (ALMA, HERSCHEL)

SAM candidate companion confirmation

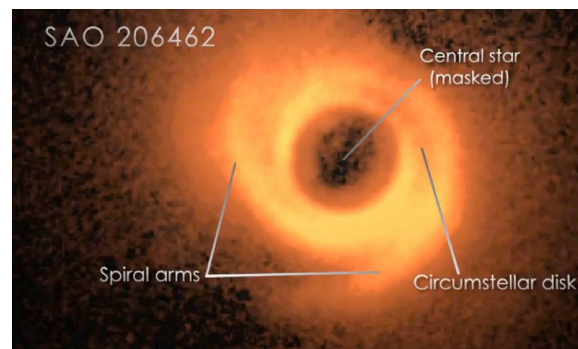
Advanced hydro. Simulations

(2 planets ? 1 binary + 1 planet ?)

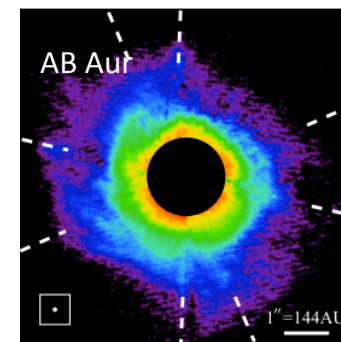
Unique case



*Rameau et al. (2012)*



*Muto et al. (2011)*



*Hashimoto et al. (2011)*